

Current Status and Plans for the uGMRT Real-time RFI Filtering System

Kaushal D. Buch, Ruta Kale, Ajith Kumar B.

Digital Backend Group

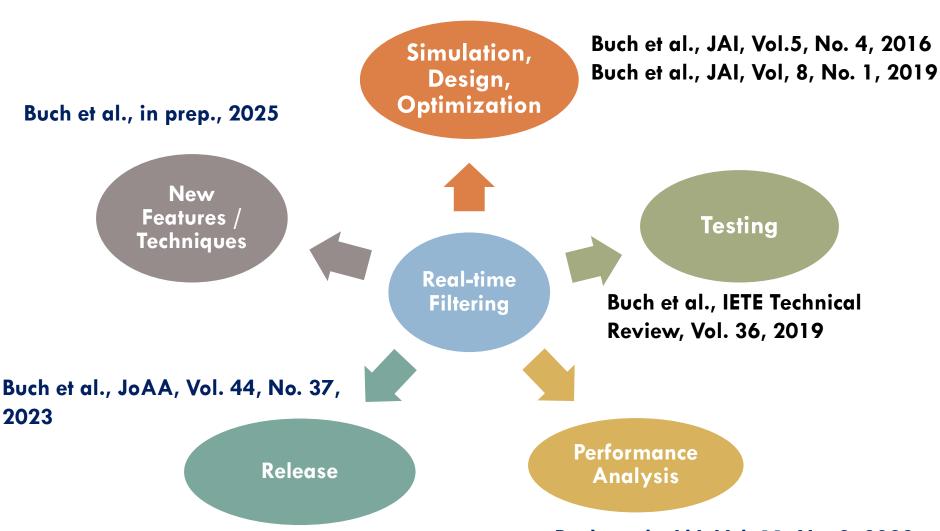
Giant Metrewave Radio Telescope (GMRT), NCRA-TIFR,

Pune, India

kdbuch@gmrt.ncra.tifr.res.in

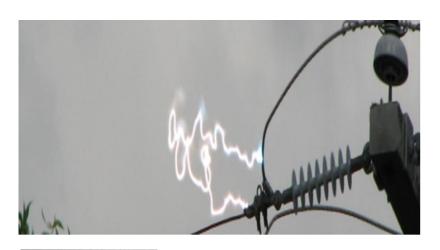
Radio Frequency Interference (RFI) 2025
Online, 22nd October 2025

RFI Filtering Development at uGMRT: Various Activities



Buch et al., JAI, Vol. 11, No. 2, 2022

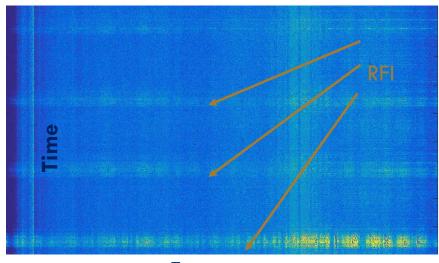
Sources of Broadband RFI











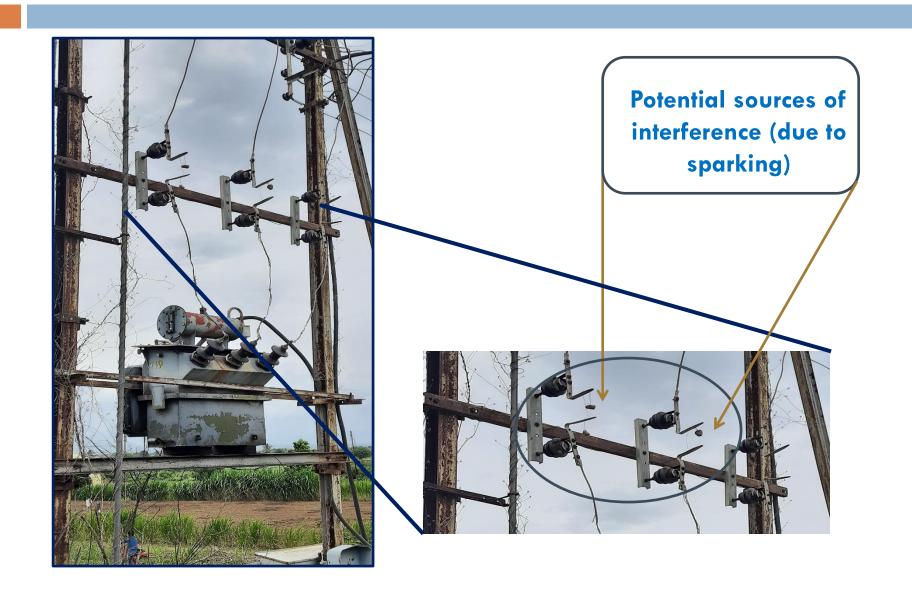
Frequency

Image Courtesy: Internet

Stray Wire on 440 kV HT Line

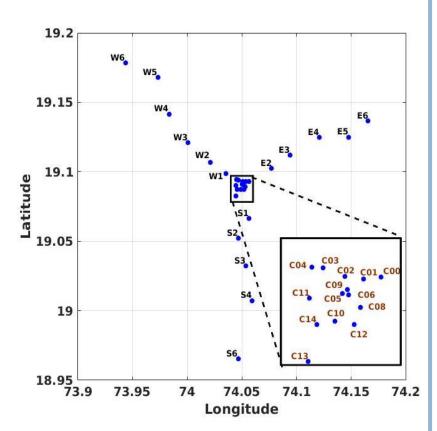


RFI from 11kV Transformers



Major sources of powerline RFI

GMRT Array



Potential Sources of Powerline RFI around Central Square and South-arm of GMRT

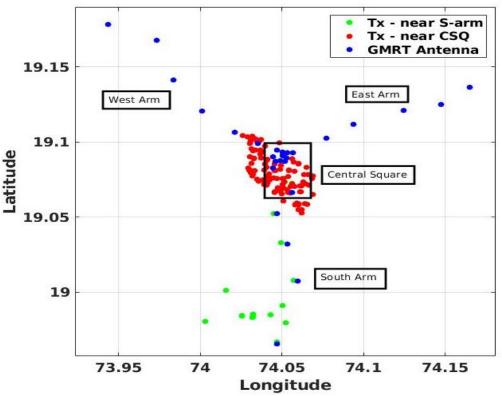
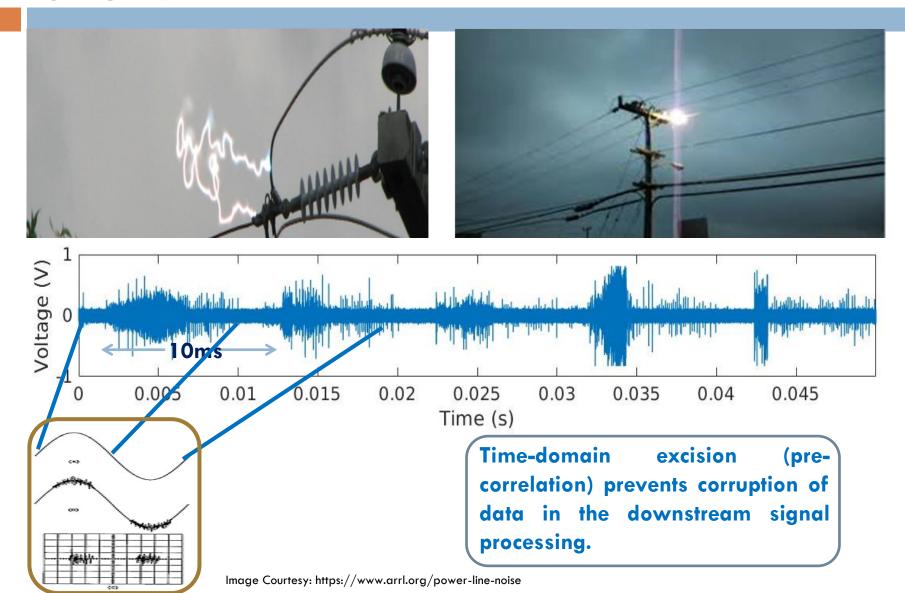


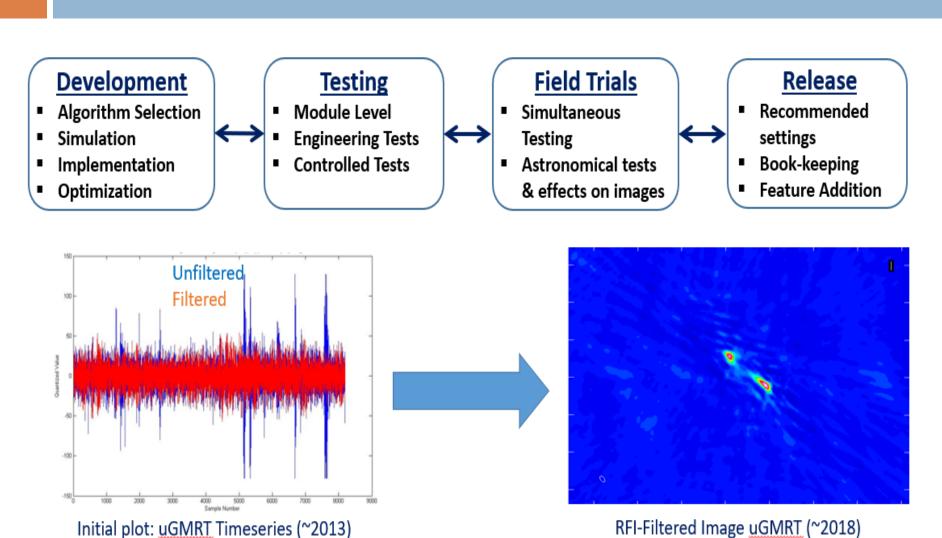
Image Courtesy: Bela Dixit

Data Courtesy: Pravin Raybole, RFI Group, GMRT

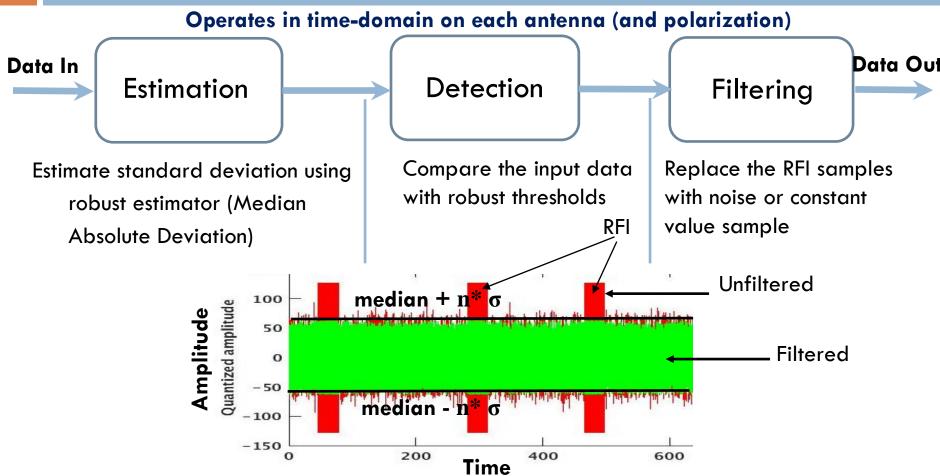
Powerline RFI at GMRT – Timeseries at 325 MHz



uGMRT RFI Filtering System: Development Flow



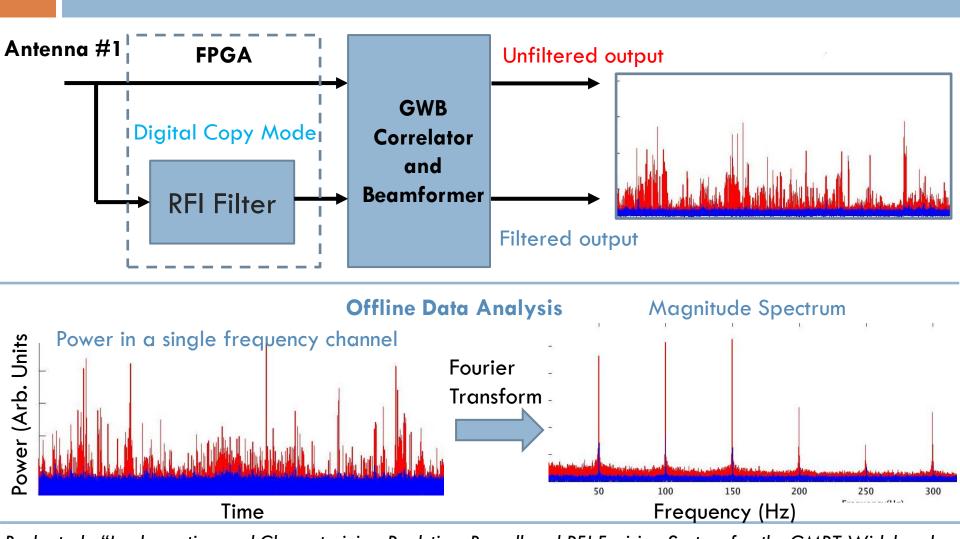
The Filtering Algorithm – Detecting and Removing Outliers



Computing recursive median in real-time was the most challenging aspect of this system

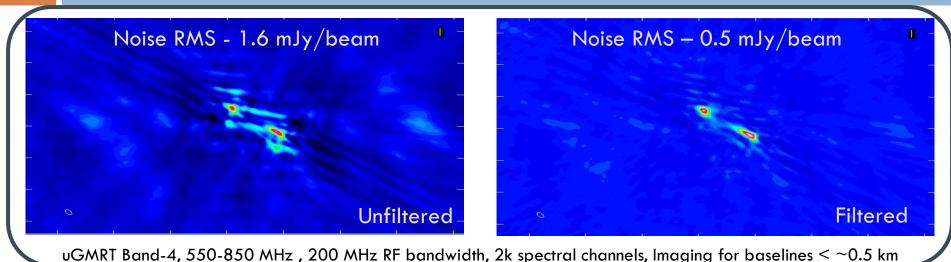
Buch et. al, "Real-time MAD-based RFI Excision on FPGA", JAI Special Issue on Interference Mitigation in Radio Astronomy, Vol. 8, No. 1, January 2019

Simultaneous Testing: Digital Copy Mode



Buch et al., "Implementing and Characterizing Real-time Broadband RFI Excision System for the GMRT Wideband Backend", IETE Technical Review, Taylor and Francis, May-June 2019

Example Results from the released system (Simultaneous Testing)



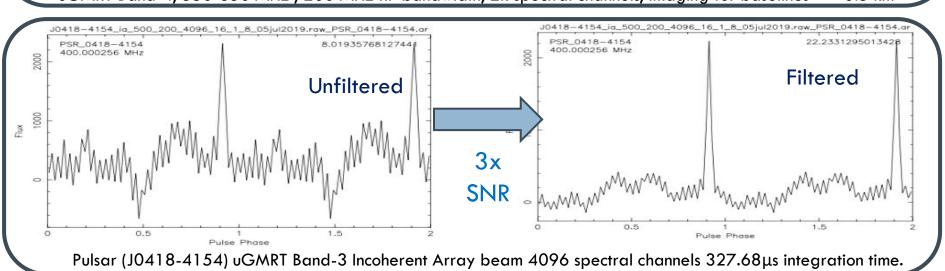
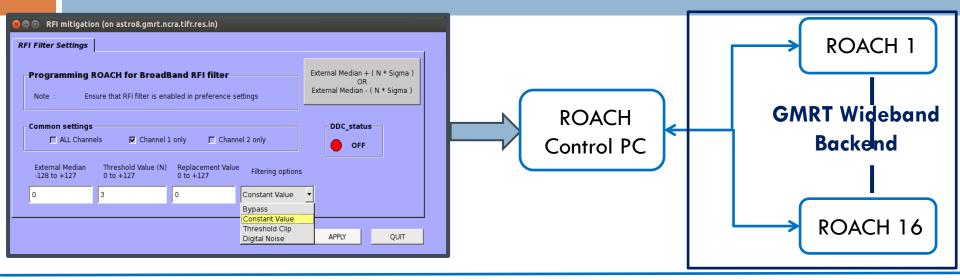
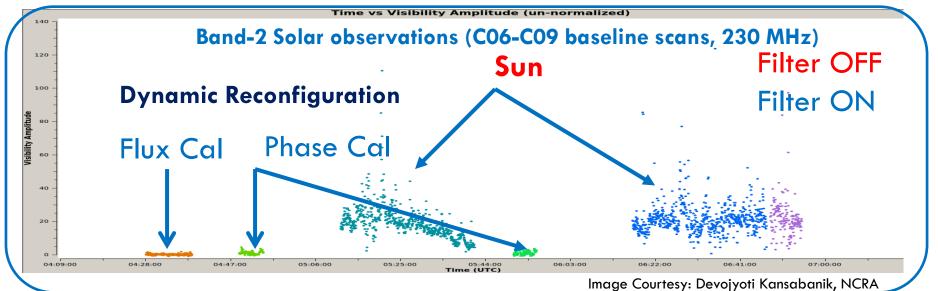


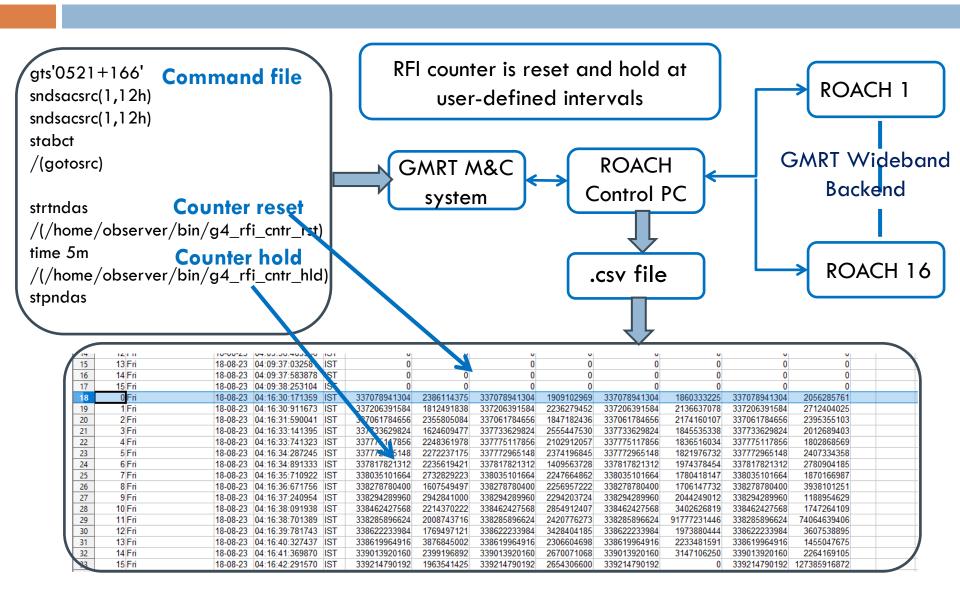
Image Courtesy: Ruta Kale, Bhaswati Bhattacharyya, Jayanta Roy, NCRA

Commissioning in the GMRT Wideband Backend (GWB)



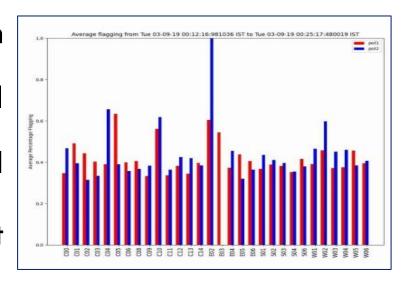


Book-keeping



Operational System at GMRT

- Full release (April 2022) with facilities like:
 - Book-keeping through fractional flagging count
 - Configurable threshold and replacement options
- Typical flagging rate of 0.5-5% at 3σ threshold
- Recommended options:
 - 3σ (pulsar), 4σ (continuum)
 - Replacement by digital noise
 - Recording counter every five minutes



Cycle 47: \sim 750 hours (\sim 35-40% of time) with RFI filtering: Continuum, Pulsar, FRB, Sun observations

Porting the system on other platforms: Ongoing development

CPU or GPU platform (GMRT Wideband Backend)

FPGA

Proposed Implementation

CPU

GPU

Tile Processing Module - Low Frequency Aperture Array SKA

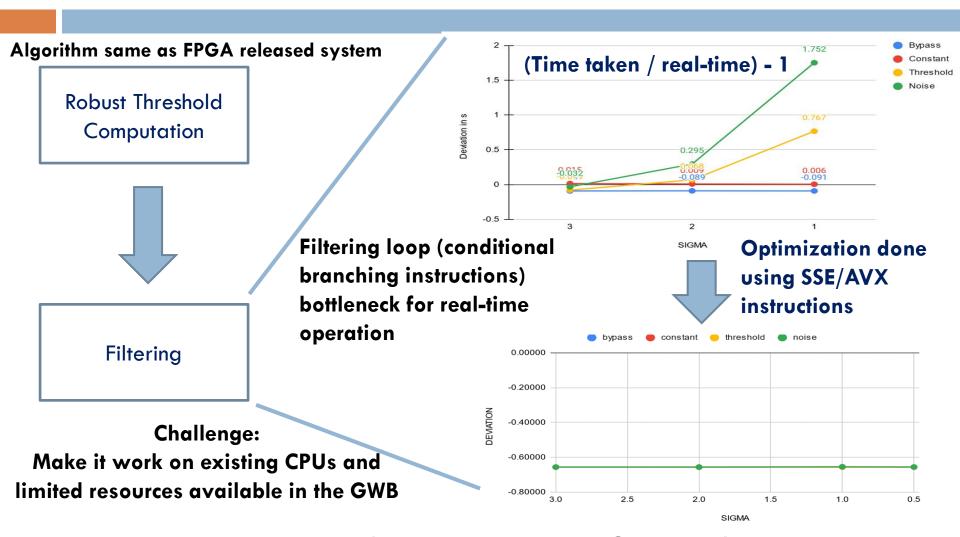
ADC Buffer

RFI Filter

Channelizer and Beamformer

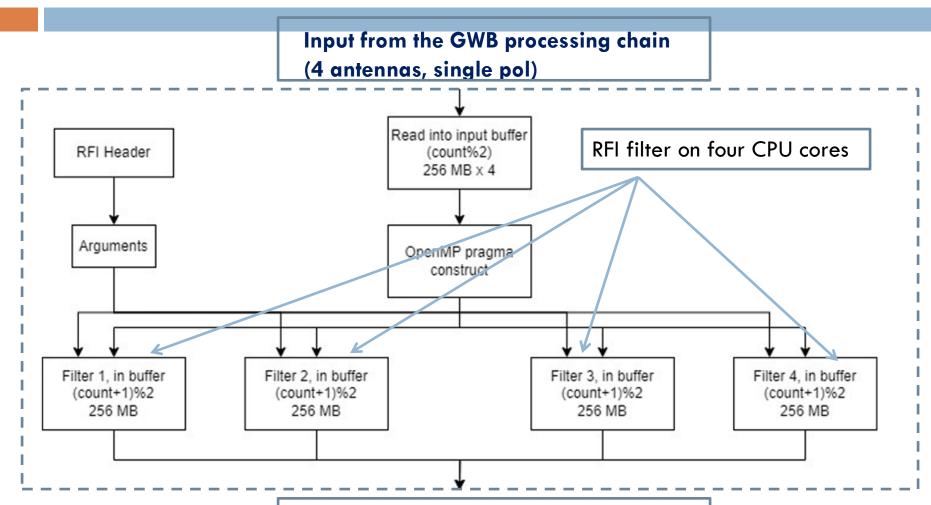
AMD-Xilinx Kintex Ultrascale+ FPGA (Implementation using VHDL)

CPU Implementation: Challenges



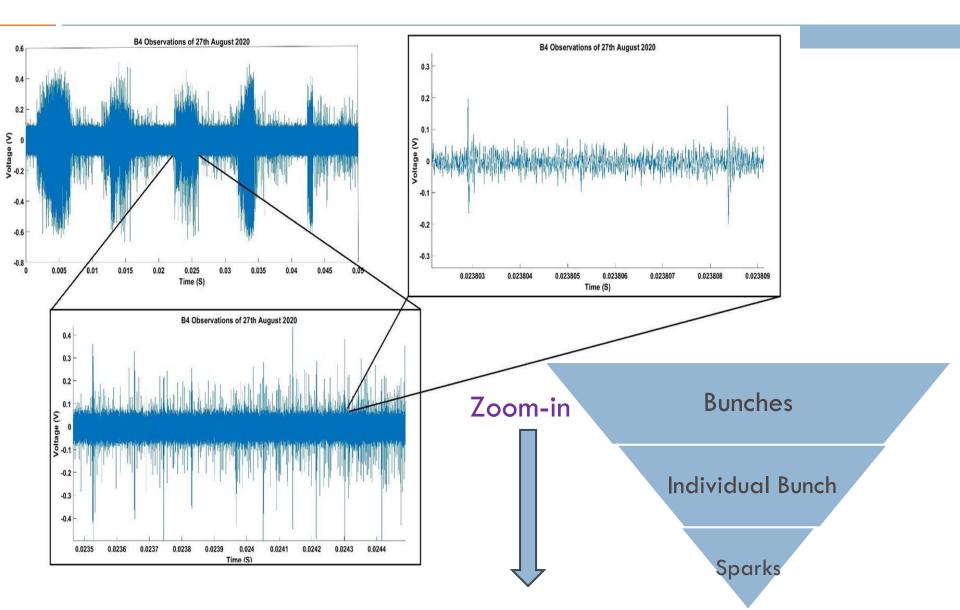
Four antennas processed on one server on CPU, each on one core

CPU Implementation: Towards Release

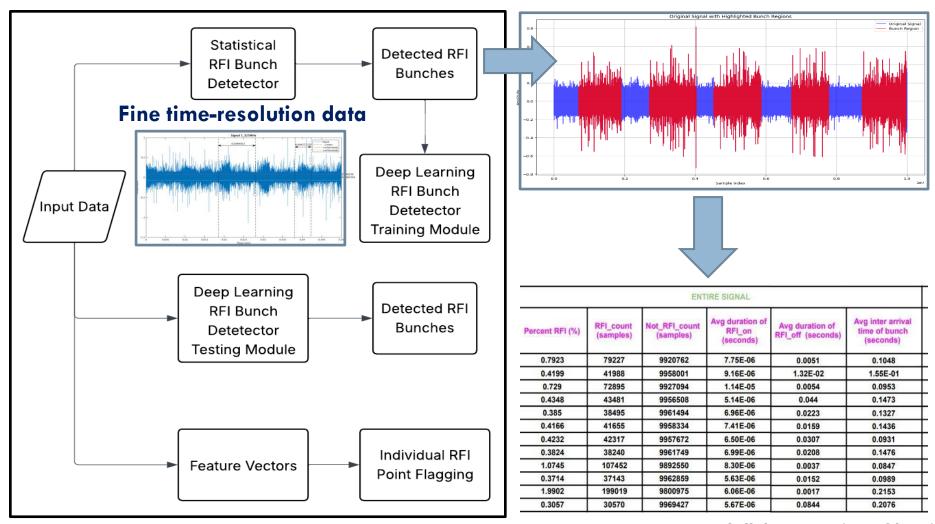


Filtered output from the GWB processing chain (4 antennas, single pol)

Powerline RFI: Composite Signal

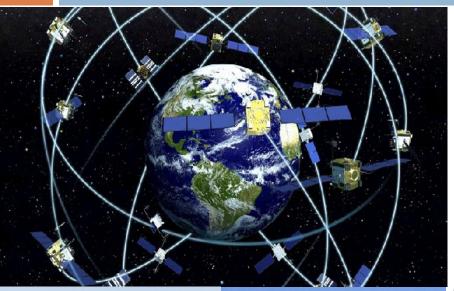


Towards Statistical Learning Approaches



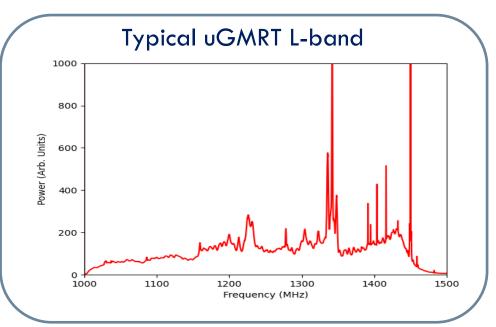
Collaborator: Divya Oberoi

Narrowband RFI

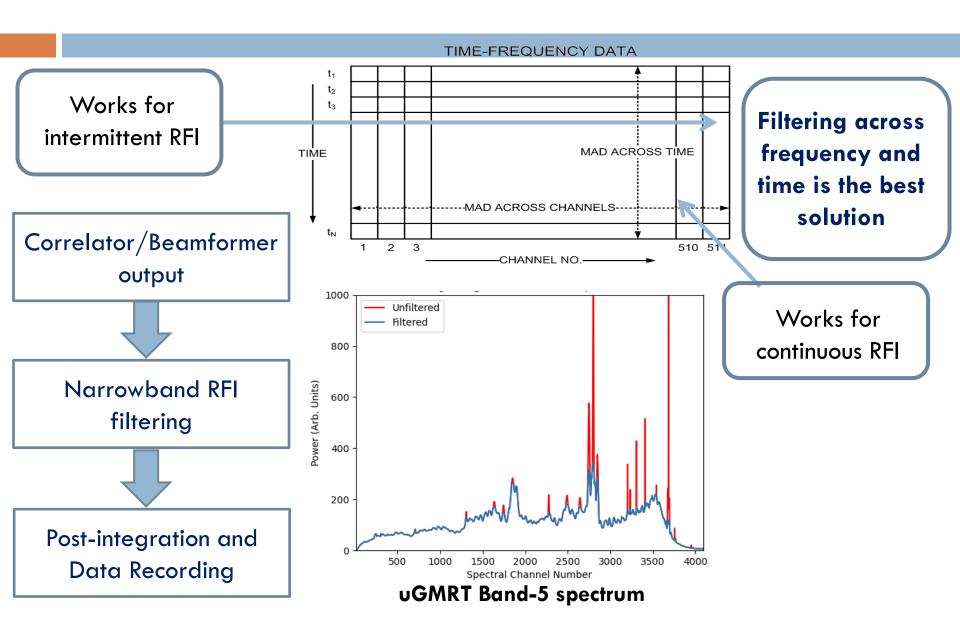


- ☐ Spectrally (and at spatially) confined
- ☐ Communication transmitters terrestrial and extra-terrestrial
- ☐ Stronger than the signal of interest (noise)

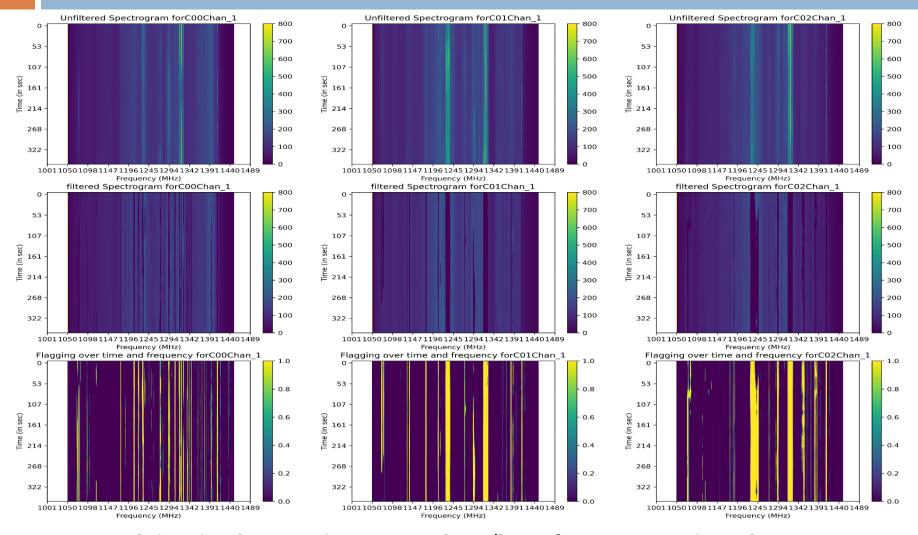




Time-Frequency Filtering



Results from offline filtering on correlator output



Real-time implementation on correlator/beamformer output is underway